



ORIGINAL PAPER

PRELIMINARY ASSESSMENT OF THE EFFECT OF NOISE ON VELOCITY
UNCERTAINTY ON THE NIGERIAN PERMANENT GNSS NETWORKSwafiyudeen BAWA ^{1)*}, Lazarus Mustapha OJIGI ¹⁾,
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ABSTRACT

With the evolution of GNSS technology, geodynamic activities can appropriately be modelled nowadays. GNSS derived time series from which velocities and their uncertainties are derived, are vital derivatives in geodynamic modelling processes. Therefore, understanding all the stochastic properties is crucial. Assuming that GNSS coordinate time series is characterized by only white noise may lead to underestimation of velocity uncertainties. In this contribution, noise behaviour of NigNET tracking stations position time series was examined by adopting WN, FL+WN, WN+RW, WN+PL. Using the maximum likelihood estimate (MLE), Akaike Information Criteria (AIC) and Bayesian Information Criteria (BIC) the quality of stochastic model or the goodness of fit of a stochastic model is determined. The results of this study show that the combination of white plus flicker noise is the best model for describing the stochastic part of NigNET tracking stations position time series.

INTRODUCTION

Generally, geodetic techniques that are satellite-based are useful tools for glacial isostatic adjustment (Lidberg et al., 2007), tropospheric modelling for numerical weather assimilation (Isioye et al., 2016), crustal motion as a result of earthquakes and tectonic strain rate (Hackl et al., 2011). For modelling crustal motion and strain rate localization, velocities and uncertainties of permanent geodetic monuments derived from Global Navigation Satellite System (GNSS) time series are often utilized (Nikolaidis, 2002; Hackl et al., 2009). Therefore, since velocities are derivatives of repeated GNSS measurement, unbiased velocities and their uncertainties are essential such that proper analysis of position time series can be attained and large variety of errors are reduced (Amiri-Simkooei et al., 2007; Hackl et al., 2011; Nistor and Buda, 2016).

Over the years, permanent GNSS stations (often referred to as Continuously Operating Reference Stations (CORS)) providing spatial and temporal information, have been set up around the world for reference frame realization and model earth geodynamics. However, these dynamics can only be modelled to some reasonable degree and possibly mitigate some errors with correct stochastic and functional models (He et al., 2017). The stochastic

part mostly considered as observational noise can be described by noise models (Goudarzi et al., 2015). Most of the stochastic parts of GNSS time series are time correlated. Mis-modelled satellite orbits, mis-modelled atmospheric effects, mis-modelled antenna phase centre effect among others are time correlated (Klos et al., 2014; Langbein and Svarc, 2019; Mao et al., 1999).

In Nigeria, an activity to set up a CORS called NIGERIAN Reference GNSS NETWORK (NigNET) which is a system of Continuous GNSS stations (see Fig. 1) commenced in 2008 by Nigeria Office of the Surveyor General of the Federation (OSGoF). The activity was gone for adding to the African Reference Frame (AFREF) and fill in as an essential fiducial system that will characterize and appear another reference outline dependent on space geodetic strategy (Jatau et al., 2010; Bawa et al., 2019).

Often, what is actually measured in a large crustal entity is the motion of geodetic monuments, from which inferential velocity and strain rate localization are done (Mao et al., 1999). With the recent reports of earth tremors (Vanguard News, 2018) and numerous proposal for a new reference frame base on space geodetic technique (Dodo et al., 2011), spatial and temporal analysis of noise on the Nigeria Permanent GNSS stations is yet to be visited



Fig. 1 Spatial distribution of NigNET tracking stations.

as done by (Klos et al., 2014; He et al., 2016; Xu and Yue, 2017)

Klos et al. (2014) who analysed more than 40 stations belonging to the ASG-EUPOS and EPN networks with 5 years of observations from the area of Sudeten, concluded that the WN+PL noise best describes the error sources for most of the analysed stations. Elsewhere, Goudarzi et al. (2015) analysed the behaviour of noise in 112 continuously operating GPS (CGPS) position time series in the eastern part of North America and found out that WN+FN is the best model that describes the stochastic part of the position time series. Xu and Yue (2017) in their assessment of the noise characteristics of daily position time series from 12 International GNSS Service sites located in China concluded that the noise model of most sites can be characterized by a combination of WN+FN.

Therefore, for geophysical studies, it is likely that a variety of time-correlated processes might dominate the error budget. More so, GNSS velocity uncertainties are usually underestimated by factors from 5 to 11 if natural only white noise that is not time correlated is considered as the dominant error source in GNSS (Hackl et al., 2011; Mao et al., 1999; Williams, 2003; Zhang et al., 1997).

To objectively select the best noise model, He et al. (2019) have investigated various criteria such as the log-likelihood value, the Akaike information criterion (AIC) and the Bayesian information criterion (BIC).

The sole aim of this paper is to identify the optimal noise model in the stochastic domain that best describes the NigNET tracking stations. Since more than half of the analyzed time series are less than 5 years (large data gap) on the average and less than 10yrs of data (short data span) (Langbein and Svarc, 2019) the study is considered preliminary.

FUNCTIONAL AND STOCHASTIC MODELS OF NIGNET TIME SERIES

FUNCTIONAL MODEL OF TIME SERIES

The functional model is a constituent of geophysical phenomena present in every GNSS time series (He et al., 2017). Therefore, the daily position of each NigNET station used for this study are combined to form time series of geodetic position using Equation (1) (Goudarzi et al., 2015; Nikolaidis, 2002; Wang, 2015). The XYZ Cartesian coordinate (earth-centered earth-fixed coordinate system) time series are converted to a topocentric coordinate which are cleaned and modelled in all three components (N, E, U) independently using Equation (1).

From the transformed daily coordinate, in Equation (1), velocity and position time series of each GNSS component (North, East and Up) accounting for first order parameters is given by (Goudarzi et al., 2015; Li et al., 2015; Nikolaidis, 2002).

$$\begin{bmatrix} N \\ E \\ U \end{bmatrix} = \begin{bmatrix} -\sin(\lambda)\cos(\phi) & -\sin(\lambda)\sin(\phi) & \cos(\lambda) \\ -\sin(\phi) & \cos(\phi) & 0 \\ \cos(\phi)\cos(\phi) & \cos(\lambda)\cos(\phi) & \sin(\lambda) \end{bmatrix} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix} \quad (1)$$

$$y(t_i) = a + bt_i + \sin(2\pi t_i) + d \cos(2\pi t_i) + e \sin(4\pi t_i) + f \cos(4\pi t_i) + \sum_{k=1}^n j_k H(t_i - t_{j_k}) + v_i \quad (2)$$

Where, $y(t_i)$ ($i = 1, 2, 3, \dots, N$) is the position at epoch t_i in years, a is the station position, b is the linear velocity, c and d are the annual and e and f are the semi-annual amplitudes of sine and cosine functions respectively. $\sum_{k=1}^n j_k H(t_i - t_{j_k})$ are offsets caused by earthquakes, environmental, equipment malfunction or change, or human intervention (Li et al., 2015), n is the number of offset, j_k is the magnitude change in the position time series at epoch t_{j_k} , H is the Heaviside step function. v_i signifies measurement error.

$$y = Ax + v \quad (3)$$

In Equation (3), $x = [a \ b \ c \ d \ e \ f \ j]^T$, y is the observation matrix and A is the design matrix. Thus, A can be formulated as given in Equation (4).

Equation (3) is solved using the principle of least squares to estimate the vector of estimated parameters, covariance matrix of the observations, covariance and cofactor matrices of estimated parameters and vector of post-fit residuals.

$$A = \begin{bmatrix} 1 & t_1 & \sin(2\pi t_1) & \cos(2\pi t_1) & \sin(4\pi t_1) & \cos(4\pi t_1) & H(t_1 - t_{j_1}) & \dots & H(t_1 - t_{j_n}) \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \ddots & \vdots \\ 1 & t_N & \sin(2\pi t_N) & \cos(2\pi t_N) & \sin(4\pi t_N) & \cos(4\pi t_N) & H(t_N - t_{j_1}) & \dots & H(t_N - t_{j_n}) \end{bmatrix}_{N(6+n)} \quad (4)$$

STOCHASTIC MODEL OF TIME SERIES

GNSS time series are perturbed by random and time correlated noise, this in-turn, results in inaccurate velocity and uncertainty estimate of GNSS observables (Goudarzi et al., 2015). However, removing outliers using designated threshold or criterion can't detect or remove some outliers (He et al., 2017). Therefore, so that realistic uncertainty can be assigned, maximum Likelihood estimation (Bos et al., 2008, 2013), overlapping Hadamard variance (OHVAR) (Xu and Yue, 2017), Least squares variance component estimation (LS-VCE) (Amiri-Simkooei et al., 2007) and Alan Variance (Niu et al., 2014) are often used to estimate the optimal noise characteristics in GNSS time series.

Noise analysis in GNSS time series does not reduce the noise but classifying the noise can identify the source of the noise and characterize them, hence, help increase accuracy and precision. Monument instability for example is known to follow the random-walk noise process. So to counter such effect, deep drill braced monuments are recommended for GNSS monuments (Langbein, 2008; Langbein et al., 1995; Williams, 2003; Zhang et al., 1997), this is because deep drill braced geodetic monuments exhibit less temporally correlated noise than any other monument type. The NigNET tracking stations are mostly on roof tops with the exception of station CGGT established on exposed bedrock.

Noise in GNSS position time series can be described as a power-law process (Williams et al., 2004) of the form,

$$P(f) = P_o \left(\frac{f}{f_o} \right)^{\kappa} \quad (5)$$

Where f_o and P_o are normalized constant, f is the spatial frequency and κ is the spectral index. Since GNSS position time series contain both white noise and time correlated noise, considering only white noise will lead to the underestimation of site velocity uncertainties by a factor equivalent to or more than 4 units (Yuan et al., 2008). Thus, there is the need for more robust models to model velocity uncertainties in GNSS. In this regard, numerous models have been proposed (Langbein, 2008; Williams et al., 2004; Zhang et al., 1997). Therefore, noise model in GNSS can be WN, WN+RW, WN+FN, WN+PL, WN+FN+RW, First-Order Gauss-Markov noise + white noise (FOGM+WN) (He et al., 2017; Williams et al., 2004).

In geophysical phenomena, spectral index range from -3 to +1. This can further be segregated into fractional Gaussian motion with spectral index of $-1 < \kappa < +1$ and fractional Brownian motion with spectral index of $-3 < \kappa < -1$ (Goudarzi et al., 2015; Wang, 2015). When $\kappa = 0$, the equivalent is termed white noise, when $\kappa = -1$ the equivalent is termed flicker noise, while $\kappa = -2$ is termed random-walk noise or Brownian motion (Mao et al., 1999).

White noise which is not time correlated can be reduced through repeated observations and averaging, is mostly caused by GNSS measurement error and hardware noise (Goudarzi et al., 2015; Mao et al., 1999). Flicker noise is regionally uniform and most common in dynamic process e.g. wobble motion of the earth, sunspot variability, uncertainties in time measure by atomic clock and undersea currents (Nikolaidis, 2002), mis-modelled of satellite antenna phase center, satellite vehicle orbits, Large scale atmospheric effect (Klos et al., 2015).

In this study, HECTOR software (Bos et al., 2013) was used for characterizing GNSS noise model using Maximum Likelihood Estimation given in Equation (6).

$$lik(\hat{v}, C) = \frac{1}{(2\pi)^{\frac{N}{2}} (\det C)^{\frac{1}{2}}} \cdot \exp\left(-0.5 \hat{v}^T C^{-1} \hat{v}\right) \quad (6)$$

Where \det is the determinant of a matrix, C is the covariance matrix of assume noise in data, N is the number of epochs, \hat{v} is the time series residuals vector between the data and the functional model. Equation (6) can be rewritten as Equation (7)

$$\ln[lik(\hat{v}, C)] = -0.5 \left[\ln(\det C) + \hat{v}^T C^{-1} \hat{v} + N \ln(2\pi) \right] \quad (7)$$

Adopting the methods of (Goudarzi et al., 2015; He et al., 2017; Jiang et al., 2014; Klos et al., 2015; Mao et al., 1999; Yuan et al., 2008) four classes of noise models adopted in GNSS applications, namely; WN, WN+PL, WN+RW and WN+FN are investigated. The choice of these models is based on the notion that they are the most common in GNSS error modelling. The different noise models adopted in this study can be written as the covariance matrix C (Bos et al., 2008; Klos et al., 2014; Goudarzi et al., 2015) as defined in the maximum likelihood.

To evaluate the quality of a chosen noise model or the goodness of fit of a chosen noise model, the Akaike Information Criteria (AIC) and Bayesian Information Criteria (BIC) (Akaike, 1974; Schwarz, 1978) were used (Bos et al., 2013).

EFFECT OF TIME CORRELATED NOISE ON NigNET VELOCITY

Zhang et al. (1997), Mao et al. (1999), Williams (2003), Bos et al. (2008), Li and Shen (2012) provided rate uncertainties for different noise models. Thus, following and adopting Li and Shen (2012), velocity uncertainty of white noise and flicker noise model is calculated as given in the Equations (8) and (9) respectively.

$$\sigma_w = \sqrt{\frac{12A_w^2}{\Delta t^2 m(m^2 - 1)}} \cong \frac{2A_w}{T} \sqrt{\frac{3}{m}} \quad (8)$$

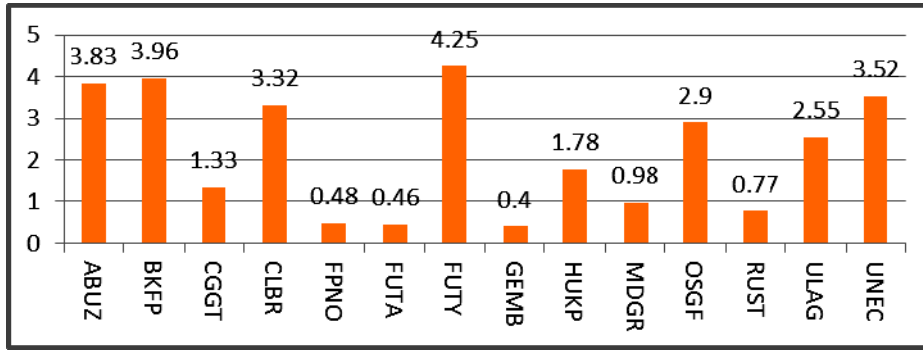


Fig. 2 NigNET data span.

Table 1 Number of Times Each Noise Model. Combination has the Lowest AIC/BIC.

Noise model	East(E)		North(N)		Up(U)		% (E,N,U)	
	AIC	BIC	AIC	BIC	AIC	BIC	AIC	BIC
WN	1	1	2	2	2	3	11.5 %	13.3 %
WN +PL	1	0	1	0	0	0	4.5 %	0 %
WN +RW	2	3	4	4	5	6	25 %	28.9 %
WN+FL	11	12	7	8	8	6	59 %	57.8 %

$$\sigma_f = \sqrt{\frac{1.78A_f^2\Delta t^{0.22}}{T^2}} \cong \frac{3A_f}{4T} \quad (9)$$

For random-walk noise, the uncertainty is given as;

$$\sigma_{rw} = \sqrt{\frac{A_{rw}^2}{\Delta t(m-1)}} = A_{rw}\sqrt{\frac{1}{T}} \quad (10)$$

Where T is the total observation span in years, Δt is the sampling interval, m is the number of observation A_w , A_f and A_{rw} are the amplitudes of white, flicker and random Walk noise respectively.

PRESENTATION OF RESULTS AND DISCUSSIONS

TIME SERIES OF NIGNET FROM FUNCTIONAL MODEL

For time series analysis, NigNET data spanning from 01/01/2011 to 12/31/2015 were processed using GAMIT/GLOBK. Percentage number of epoch is presented in Figure 2. Co-seismic displacement, power outage, instrument change, direction of motion of station etc., are typical information extractable from GNSS time series.

The time series of all the stations showed no evidence of co-seismic, post-seismic nor inter-seismic displacement in the study periods. Large data gaps in the time series are due to power outage and possibly antenna mal-function and change as was seen in station log file of station CGGN.

ASSESSMENT OF STOCHASTIC MODEL OF NigNET TIME SERIES

As highlighted earlier, the stochastic models WN, WN+PL, WN+RW, and WN+FL in the form of noise were considered in analysing the geophysical parameters. A threshold of three time inter-quartile range (3IQR) as highlighted earlier is used to remove outliers. Furthermore, the relative goodness of fit of each noise was tested using AIC/BIC. The model with the least AIC/BIC is considered the optimal model that describes each time series component.

From Table 1 and Figure 3, it is observed that the optimal noise model that describes the source of noise in estimated geophysical parameter on the NigNET is the WN+FL noise model because the percentage of a number of times WN+FL noise model has the lowest AIC is 59 % and 57.8 % in the BIC respectively. This is consistent with the studies of He et al. (2016, 2017), Mao et al. (1999), Wang (2015), Williams et al. (2004) and Zhang et al. (1997).

Questions begin to arise when AIC and BIC are not in agreement as presented in the Table 1 and Figure 3. It is observed that RW+WN (25 %/28.9 %) is the next choice of model followed by white (WN) (11.5 %/13.3 %) AIC/BIC. The results clearly depicts why it is very difficult to consider just one noise model to characterize the error sources in a GNSS station. However, there is a small mean discrepancy between AIC/BIC (1.25 ~2.85 %) over all stations for the North East and Up components. These values are the reasons why questions arise as to the choice of noise model. An alternate method is the use of visual intuition on the fit of geophysical signal estimate on the observed GNSS time series to make decision (He

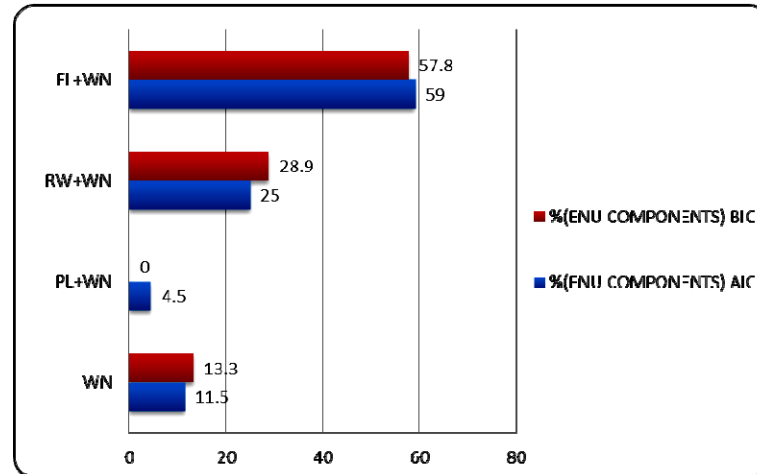


Fig. 3 Bar graph of number of times each noise model combination has the lowest AIC/BIC.

Table 2 Amplitude of white and flicker noise models.

Stations			East		North		Up	
	Lat	Long	White	Flicker	White	Flicker	White	Flicker
ABUZ	11.15174	7.64869	1.559	3.522	0.988	2.632	3.207	9.746
BKFP	12.46858	4.22924	1.462	3.506	1.012	2.756	3.064	9.464
CGGT	10.1231	9.11831	1.417	3.875	0.968	2.940	3.741	5.989
CLBR	4.9503	8.35157	2.429	3.794	1.669	3.041	6.038	11.316
FPNO	5.43457	7.03324	2.132	0	1.458	3.231	6.18	0
FUTA	7.29864	5.13644	2.748	0	1.554	2.011	5.448	7.174
FUTY	9.34974	12.49780	1.793	3.899	1.166	2.966	3.65	11.215
GEMB	6.9172	11.18394	1.972	0	1.099	0	4.203	5.44
HUKP	12.92115	7.59091	1.975	3.236	1.208	3.098	3.616	9.484
MDGR	11.83809	13.1310	1.499	3.176	0.965	3.358	3.266	7.739
OSGF	9.02767	7.48634	1.849	3.361	0.836	2.285	3.976	10.336
RUST	4.80184	6.97852	1.624	4.929	1.379	3.049	6.027	0
ULAG	6.51733	3.39762	1.937	3.037	1.218	2.762	3.994	10.985
UNEC	6.42481	7.50499	1.848	4.299	1.257	3.293	4.436	10.344

et al., 2017). Figure 4 is a typical example of spectral plots of stations ABUZ and BKFP of the NigNET tracking stations. The plots show how well each noise model fits into the observed time series. As can be seen in the spectral plot of station ABUZ for example, flicker WN+FN optimally describes it, but WN+RW best describes the noise in the vertical component.

SPATIAL DEPENDENCE OF NigNET NOISE AMPLITUDE

The question as to whether there is spatial dependence of amplitudes in a global case as pointed out by Mao et al. (1999) and Williams et al. (2004) is applicable to the NigNET case is significant. Interestingly, as presented in Figure 5, there is a significant spatial dependence between the amplitude of white noise in the Up component which corroborates the finding of Mao et al. (1999) and Williams et al. (2004). As for the North and East components, the dependence of amplitude on these components is not obvious. Furthermore, Figure 6 shows that FN has no convincing significant variation in latitude (see also Table 2), hence corroborates the findings of Williams et al. (2004). Therefore, the finding in the present study is conclusively in

conformity with the global generalization of Mao et al. (1999) and Williams et al. (2004). Note that the East, North and Up labels in Figures 5 and 6 are East amplitude, North amplitude and Up amplitude respectively.

EFFECT OF TIME CORRELATED NOISE ON NigNET VELOCITY

To further investigate the influence of the estimated geophysical parameters, the velocity of individual stations and their uncertainties considering the various noise models adopted in this study were compared. Considering only white noise as the error source in GNSS time series could result to underestimation of velocity errors by a factor of 5-11 mm/yr (Mao et al., 1999). This is analogous to Bos et al. (2008), where it was justified that considering only flicker noise could also result to underestimation of velocity uncertainty by factor of 6-13mm/yr.

In this study, Equations (8), (9) and (10) are used to compute the velocity uncertainties as presented in Table 3 and subsequently used to plot the velocity vectors as presented in Figures 7, 8 and 9 for WN, WN+RW, and WN+FN model respectively.

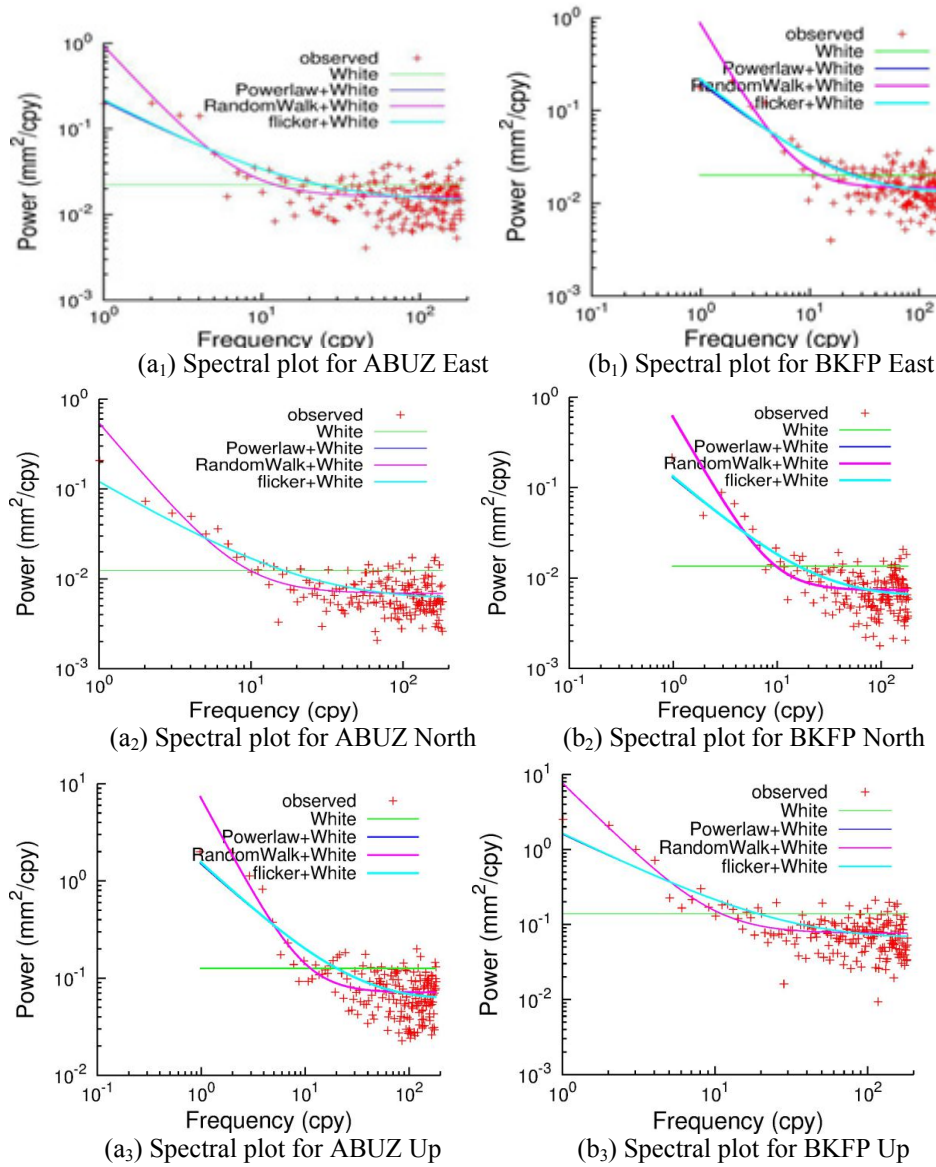


Fig. 4 Power spectral density (PSD) of residual time-series of stations ABUZ and BKFP in the North, East and Up component, respectively.

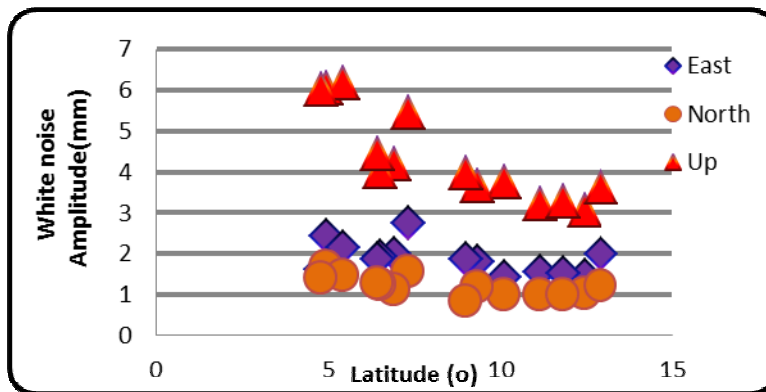


Fig. 5 Latitude dependence of white noise amplitude.

When only white noise is considered, model results to underestimation of velocity uncertainty by a factor of 7 mm/yr, 6 mm/yr and 6 mm/yr on the average for the East, North and Up components respectively for the case of WN+RW noise. More so, assuming only WN model results to underestimation

of velocity uncertainty by a factor 4 mm/yr, 5 mm/yr and 4 mm/yr on the average for the East, North and Up component respectively for the case of WN+FN. This corroborates the findings of Mao et al. (1999) and Bos et al. (2008).

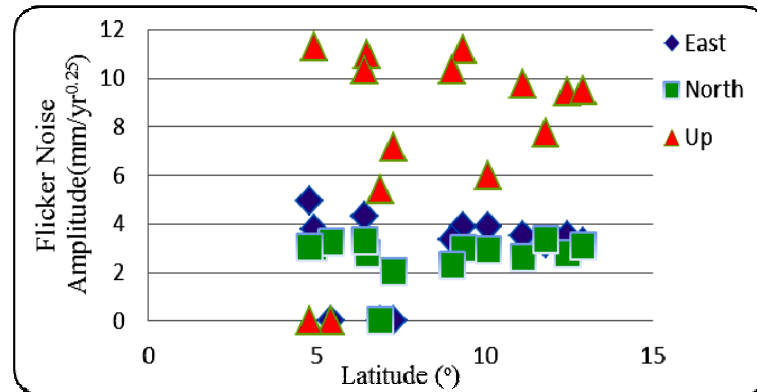


Fig. 6 Latitude dependence of flicker noise amplitude.

Table 3 Velocity uncertainties from different noise model.

Station	Long	Lat	WH			WH+RW			WH+FL			Data Span (yrs)
			E	N	U	E	N	U	E	N	U	
ABUZ	7.649	11.152	0.05	0.036	0.051	2.190	1.691	6.592	0.727	0.539	1.986	3.83
BKFP	4.229	12.469	0.044	0.036	0.044	2.070	1.737	5.975	0.698	0.545	1.863	3.96
CGGT	9.118	10.123	0.257	0.203	0.257	4.185	2.338	6.174	2.353	1.772	3.819	1.33
CLBR	8.352	4.950	0.082	0.061	0.082	2.099	1.607	6.895	0.930	0.737	2.737	3.32
FPNO	7.033	5.435	1.166	0.962	1.166	1.166	3.648	3.381	1.166	5.846	3.381	0.48
FUTA	5.136	7.299	1.606	0.979	1.606	1.606	0.979	3.442	1.606	4.187	14.881	0.46
FUTY	12.498	9.350	0.047	0.034	0.047	2.164	1.666	7.207	0.725	0.548	2.055	4.25
GEMB	11.184	6.917	0.926	0.785	0.926	1.409	0.785	3.252	1.409	0.785	13.202	0.4
HUKP	7.591	12.921	0.274	0.200	0.274	2.884	2.597	8.927	1.607	1.454	4.442	1.78
MDGR	13.131	11.838	0.341	0.286	0.341	4.064	3.661	8.613	2.711	2.750	6.534	0.98
OSGF	7.486	9.028	0.081	0.055	0.081	1.956	1.319	6.975	0.937	0.622	2.819	2.9
RUST	6.979	4.802	0.597	0.467	0.597	6.937	2.936	1.601	5.232	3.336	1.601	0.77
ULAG	3.398	6.517	0.097	0.071	0.097	1.931	2.247	7.676	0.979	0.867	3.409	2.55
UNEC	7.505	6.425	0.066	0.054	0.066	2.753	2.035	6.284	0.967	0.736	2.326	3.52

Even though WN+FN is the optimum noise model to characterize error source in NigNET tracking stations, from Table 3, the largest uncertainties are the most obvious from the combination of WN+FN. Similarly, Figures 10, 11 and 12 are bar graphs depicting compared uncertainties from the choice of the noise models as computed from Equations (8), (9) and (10). In Figure 11, the velocity uncertainty for stations FPNO and FUTA are the same for the entire models. More so, in Figure 11, the station FUTA shows that the uncertainties for WN and WN+RW are alike. Also, station GEMB has same velocity uncertainty for the noise model. Furthermore, the uncertainties when WN+FN and WN+RW models are the same for station FPNO.

SUMMARY

GNSS derived time series from which velocity and their uncertainties are derived an important derivatives in geophysical applications such as crustal motion plate among others. Understanding the stochastic part is essential so that realistic uncertainty can be attained. This paper identifies the optimal stochastic noise model best describes the NigNET

tracking stations. The results of this study show that the combination of W+FN is the optimal noise model that describes the stochastic part of NigNET tracking stations position time series.

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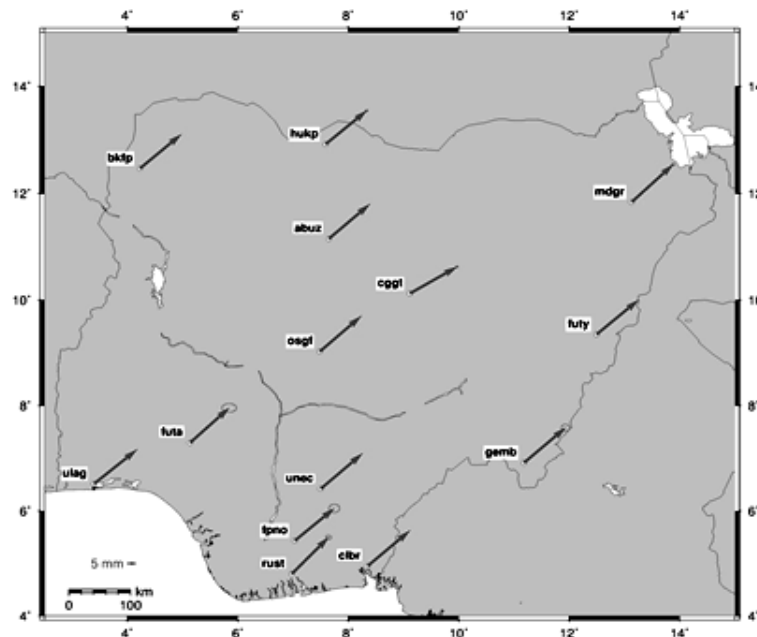


Fig. 7 Velocity field of NigNET stations and their uncertainty for white noise model only at 95 % confidence interval.

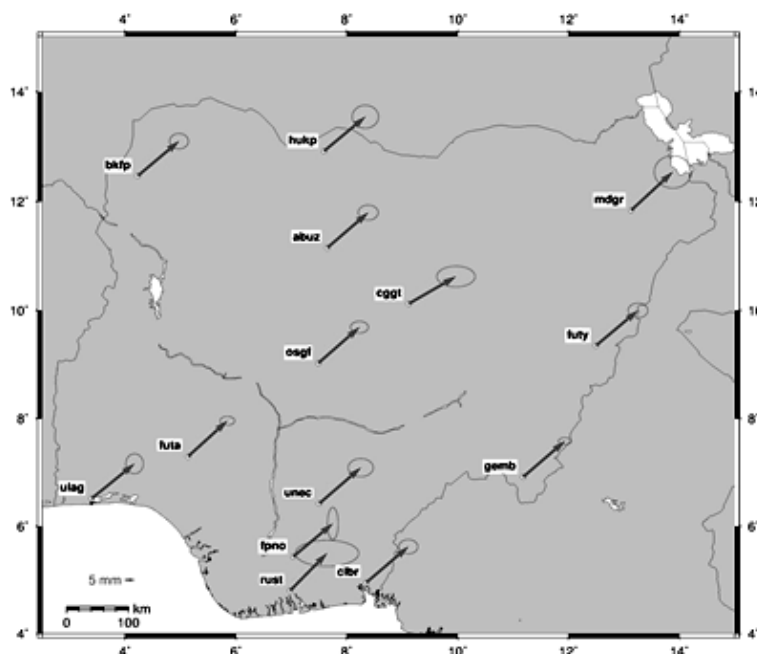


Fig. 8 Velocity field of NigNET stations and their uncertainty for white plus randomwalk noise model only.

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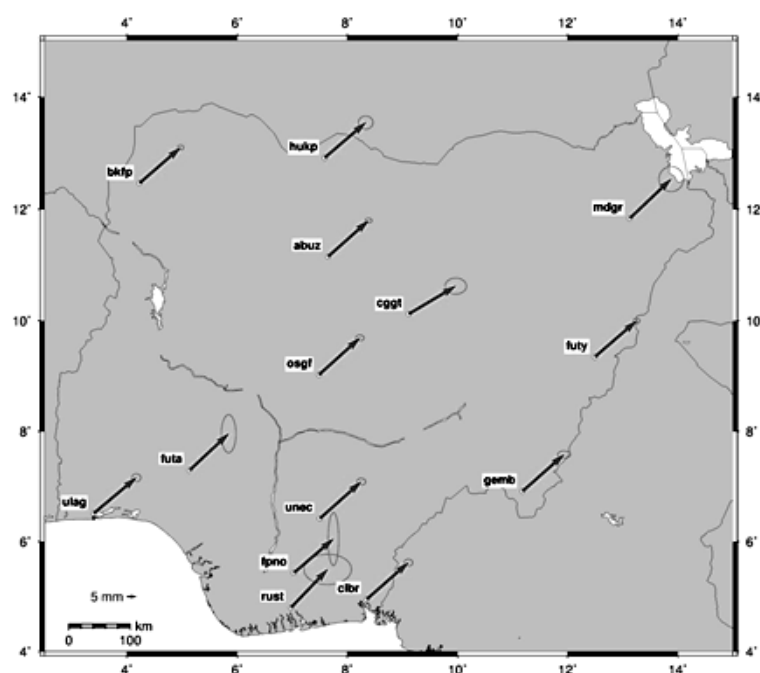


Fig. 9 Velocity field of NigNET stations and their uncertainty for white plus flicker noise model only.

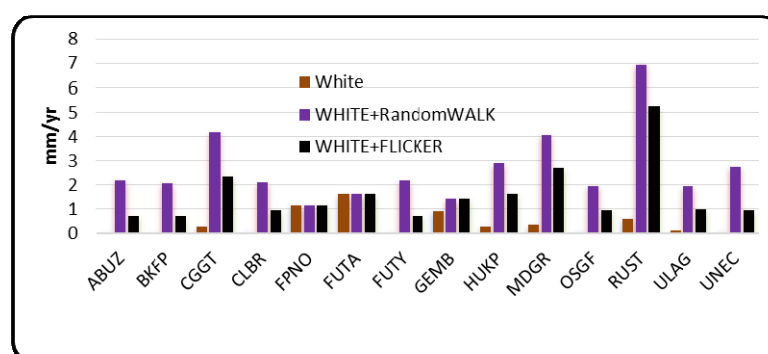


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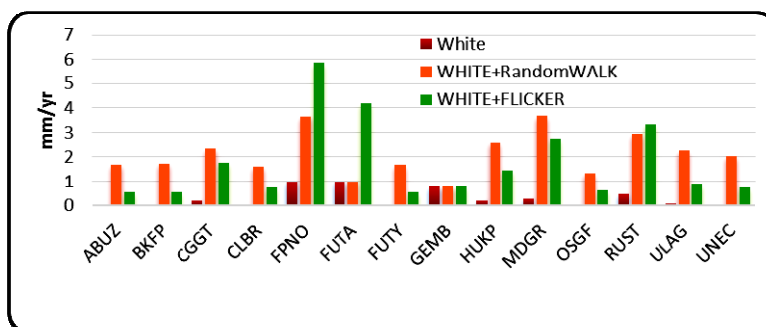


Fig. 11 Comparison of velocity uncertainty in north component from the combination of noise models.

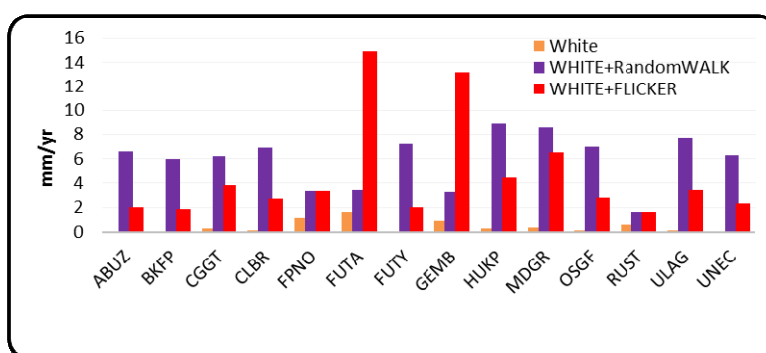


Fig. 12 Comparison of velocity uncertainty in Up component from the combination of noise models.

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